An Object-Oriented Simulation Program for CMS

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Authors:
Outline

- OSCAR overview
- Interfaces and services
- CMS detector simulation and validation
  - Magnetic field
  - Tracker
  - ECAL
  - HCAL
  - Muons
  - Forward detectors
- Parameterized simulations
- Heavy ion simulation
- Production
- Summary and outlook
OSCAR overview

Object Oriented Simulation for CMS Analysis and Reconstruction

- Full CMS simulation based on the Geant4 toolkit
- **Geant4**: physics processes describing in detail electro-magnetic and hadronic interactions; tools for the CMS detector geometry implementation; interfaces for tuning and monitoring particle tracking
- **CMS framework**: application control, persistency, common services and tools (magnetic field, generator interfaces and support for MC truth, infrastructure for hits and readout units,…), “action on demand” to selectively load desired modules, configure, tune application
- CMS changed from CMSIM/GEANT3 to OSCAR/GEANT4 end 2003;
- OSCAR used for substantial fraction of DC04 production; will be used for physics TDR production
- **CPU**: OSCAR ≤ 1.5 x CMSIM - with lower production cuts!
- **Memory**: ~110 Mb/evt for pp in OSCAR ≈ 100 Mb in CMSIM
- **Robustness**: ~1/10000 crashes in pp events (mostly in hadronic physics) in DC04 production to 0 crashes in latest stress test (800K single particles, 300K full QCD events)
**Interfaces and services**

- Application steering handled by CMS framework; CMS RunManager implements functionality required for G4 running and provides handles to the G4 run, event, track and step - as required for application configuration and monitoring; manages random number and cross-section table storage and retrieval.
- Detector geometry construction automated via Detector Description Database which converts input from XML files managed by Geometry project; XML files selected by user-defined configuration.
- Generator input (via RawHepEvent CMS format and recently HepMC) converted to G4Event; specific generator type and event format (particle gun, Pythia, etc from ntuple, ASCII, database etc) run-time configurable.
- Interface from CMS magnetic field services to G4; field selection run-time configurable; propagation parameters via XML.
- Infrastructure for physics lists (run-time selection of list and process types, optional activation of γ/e-nuclear and synchrotron radiation, misc. customizations) and production cuts (the latter via XML).
- User actions (monitoring, tuning) via dispatcher-observer pattern for pointer to observable entity.
- Persistency, histogramming, monitoring etc transparently through CMS framework (COBRA).
CMS Detector

13m x 6m Solenoid: 4 Tesla Field
→ Tracking up to h ~ 2.4

22m Long, 15m Diameter

→ > 1 M geometrical volumes;
   > 12 M readout channels

Muon system in return yoke

First muon chamber just after solenoid
→ Extended lever arm for \( p_T \) measurement

ECAL & HCAL Inside solenoid

Sliced view of CMS barrel
Magnetic Field

Field Map (TOSCA calculation)

Designed to optimize simulation and reconstruction

Based on dedicated geometry of “magnetic volumes”

Decouple volume finding and interpolation within a volume

Time spent in magnetic field query (P4 2.8 GHz) for 10 minimum bias events (wit delta=1mm) 13.0 vs 23.6 s for G3/Fortran field

⇒ new field ~1.8-2 times faster than FORTRAN/G3

GEANT4 volumes can be connected to corresponding magnetic volumes ⇒ avoid volume finding ⇒ potential ~2x improvement

With G4, also possible to use local field managers for different detectors
Tracker

Detailed description of all active and passive components; material budget

Critical requirements for physics studies with tracker

Correct, navigable Monte Carlo truth (particle, track, vertex, history) with trace-ability of initial primary particle

Special treatment of hard brem with the assignment of new track for electron above threshold (500 MeV)

⇒ Extensive validation in terms of tracking and hit distributions
Hits from minimum bias events in Tracker

Pixel cut in G3 too high
⇒ 10% increase expected
±5% differences in Si not significant

Raw simulated hits
Reconstructed hits
Electromagnetic Calorimeter (ECAL)

Comparisons with CMSIM/G3 and test beam data
- Energy and position resolution, shower shape
- Hadronic showers
- Level-1 e/m trigger response
- Preshower response
- Performance studies

Energy resolution

Position resolution
Red – OSCAR_2_3_0_pre5, black – CMS132
ECAL cont’d

Single crystal containment: $E_{1x1}/E_{3x3}$ versus position

Preshower response
Hadronic Calorimeter (HCAL)

Energy resolution

Extensive validation program with comparisons to G3 and several test beam data sets, incl. combined ECAL-HCAL runs; also in context of LCG simulation physics validation project.

Non-linearity in energy response
**HCAL cont’d**

Longitudinal shower profile

\[ e/\pi: G3 \sim 3\% \text{ higher}, \ G4 \sim 4\% \text{ lower} \]

**Energy resolution**  
\[ E\text{Cal} + H\text{Cal data} \]

Longitudinal shower profile
Muons

Detector and physics validation in terms of tracking and hit distributions with single μ’s, Drell-Yan pairs (Mll=2TeV) and physics events H→ZZ→4µ

Muon detector layout

![Muon detector layout diagram](image-url)
Muons cont’d

Trigger efficiency vs $\eta$

CMSIM

H$\to$ZZ$\to$4$\mu$, $M_H=150$ GeV
Forward Detectors
for diffractive and heavy ion physics

**Castor** Calorimeter
at 14.37 m (5.3 < \(\eta\) < 6.7)
(Tungsten/quartz plates)

**Totem** Telescopes
7.5 m < \(z\) < 13.5

**ZDC**: Zero degree calorimeter
(Tungsten/quartz fiber) at 140 m

1 TeV neutron
Parameterized Simulations

G4FLASH

Implementation of fast EM shower simulation in Geant4/OSCAR, using GFLASH parameterized showers (spot density) - tuning in progress…

Timing studies

<table>
<thead>
<tr>
<th>Electron energy</th>
<th>Time/event full simulation</th>
<th>Time/event fast simulation</th>
</tr>
</thead>
<tbody>
<tr>
<td>1 GeV</td>
<td>0.8 s</td>
<td>0.5 s</td>
</tr>
<tr>
<td>10 GeV</td>
<td>1.9 s</td>
<td>0.6 s</td>
</tr>
<tr>
<td>100 GeV</td>
<td>16 s</td>
<td>0.7 s</td>
</tr>
<tr>
<td>300 GeV</td>
<td>57 s</td>
<td>1.0 s</td>
</tr>
</tbody>
</table>

Geant4 6.2 - full vs fast

50 GeV electrons
Heavy Ion Simulation

performance optimization with a twist…

G3/CMSIM: chop event in slices of 100 tracks each and run them separately; *needed due to limitations from ZEBRA*

OSCAR/Geant4: run full HI events

Factor 5 performance improvement by improved calorimeter track selection and hit processing

…effect entirely negligible in pp events!

55K generated particles, with 97K tracks from 80K vertices kept at the end of event

<table>
<thead>
<tr>
<th>Event cut in slices of 100 particles</th>
</tr>
</thead>
<tbody>
<tr>
<td>CMSIM - G3</td>
</tr>
<tr>
<td>OSCAR_2_4_5 – G4 5.2</td>
</tr>
<tr>
<td>OSCAR_3_4_0 – G4 6.2</td>
</tr>
</tbody>
</table>

Full event

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</table>

*2.3 CPU hrs on P4 3.2 GHz*
Production

OSCAR 2.4.5 in use for 10 months; longest-used version of any s/w in production; accounts for 35M of 85M events; G3 simulation ‘officially dead’

OSCAR 2.4.5 released

Wall-clock time normalized to 1GHz CPU

Peak not moved, but tail significantly narrower. Nicer for production, easier to spot stuck jobs
Summary and Outlook

In CMS, OSCAR, the OO simulation program based on the Geant4 toolkit, has successfully replaced its Fortran/Geant3 predecessor. It has been validated and adopted by all CMS detector and physics groups. It has proven robust and performant, easily extensible and configurable.

CMS has now entered sustained-mode production: 10M physics events/month through the full chain (simulation, digitization, …, DSTs) ⇒ A lot more good

**physics with OSCAR**

Higgs event \(m_H=180\) in CMS Tracker   SUSY event (leptons, missing \(E_T\))

(visualization with IGUANACMS)